

Introductory Astronomy

Week 5: Stellar Evolution

Clip 15: Neutron Stars

What is Left of Core?

- **Electron** degeneracy cannot stop collapse – few **electrons**
- **Neutron** degeneracy pressure at density
 $6 - 7 \times 10^{17} \text{ kg/m}^3$
- $1.4M_{\odot}$ in $R \sim 10 \text{ km}$
- Surface gravity $1.9 \times 10^{11} g$

- Physics is **relativistic**
- **Chandrasekhar Limit**

$$MR^3 = \text{const}$$

$M_{Ch} = 2.2 - 2.9M_{\odot}$ depends on **rotation**

- **Rapid Rotation** expected
- High **magnetic field** frozen in
 $B_{\text{ns}} \sim B_{\text{wd}} \left(\frac{R_{\text{ns}}}{R_{\text{wd}}} \right)^2$

Discovery

- Physics Predictions:

- Rapid Rotation

$$MvR \sim MR^2/P$$

$$P_{\text{ns}} = P_{\text{c}} \left(\frac{R_{\text{ns}}}{R_{\text{c}}} \right)^2 \sim 0.005 \text{ s}$$

- Intense magnetic field

$$B_{\text{ns}} \sim B_{\text{c}} \left(\frac{R_{\text{ns}}}{R_{\text{c}}} \right)^2 \sim 10^{12} B_{\odot}$$

- High Temperature

$$T \sim 10^6 \text{ K } \lambda_{\text{max}} = 2.9 \text{ nm } L \sim 0.25 L_{\odot}$$

- Bell 1967: Periodic 1.337s Radio pulses: LGM?

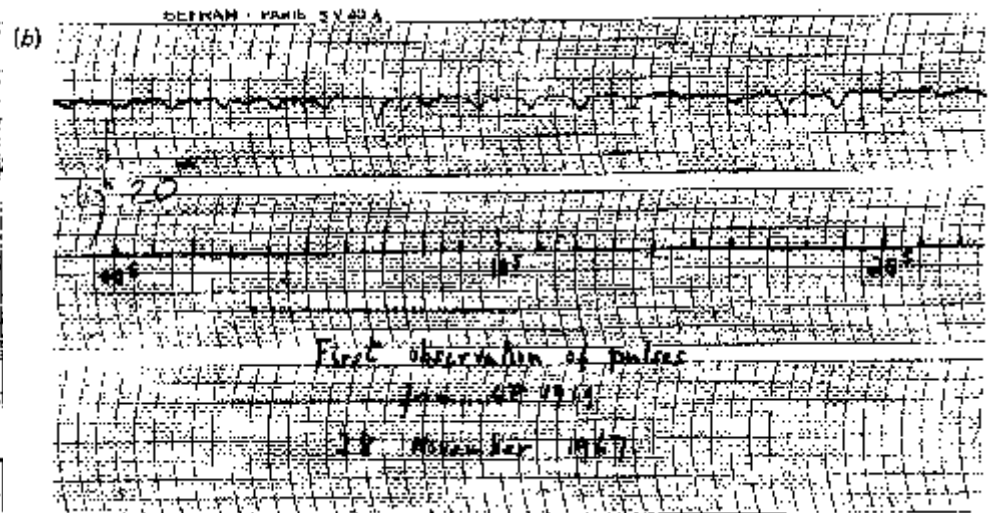
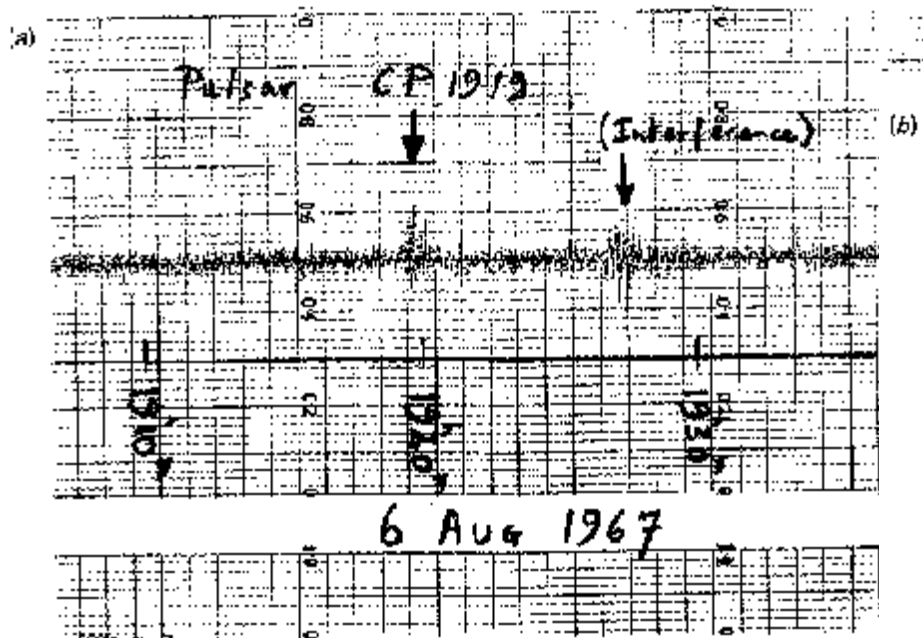
- Quickly found other sources: natural

- Soon find many pulsars

$$0.2 \text{ s} \leq P \leq 2 \text{ s}$$

Slow down in 10^7 yr

LGM Data



What are Pulsars?

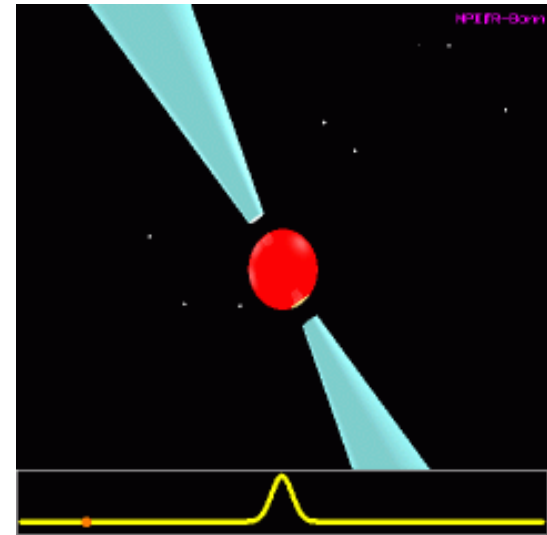
- Rotating star breaks up

$$v^2/R = 4\pi^2 R/P^2 \sim GM/R^2$$

$$P = 2\pi\sqrt{\frac{R^3}{GM}}$$

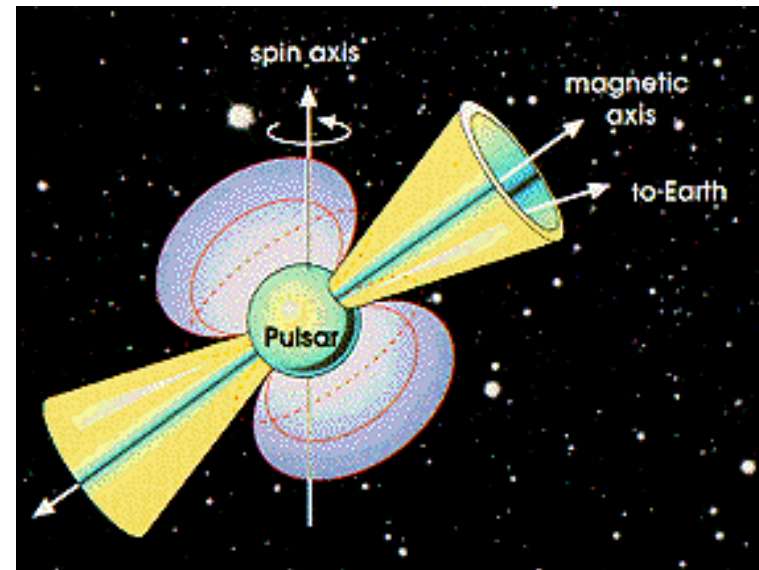
- Only NS dense enough to survive $P \sim 1\text{s}$
- Emission aligned to magnetic axis - tilted

- Crab pulsar $P = 0.0333\text{s}$:
Neutron star SN remnant



How They Work

- General Idea: Rapidly changing intense **magnetic** field creates intense **electric** field
 - Lifts **charged** particles from **polar** regions into **magnetosphere** dragged around by rotation
 - Accelerated to **relativistic** speeds – emit **synchrotron** radiation at all wavelengths in direction of magnetic **axis**
 - Emitted **energy** slows **rotation**
 - **Luminosity** of **Crab nebula** agrees with observed rate of slowing of **pulsar**
- Pulsars observed in **all bands**



Credits

- CP1919 Data, Animation:
<http://pulsar.ca.astro.it/pulsar/PSR.html>
- Pulsar Model: NASA/GFC
http://imagine.gsfc.nasa.gov/docs/science/known_1/pulsars.html