In the first year of the period Chih-ho, the fifth moon, the day chi-ch'ou, a guest star appeared approximately several [degrees] southeast of Thien-kuan. After more than a year it gradually became invisible. (1054)

Introductory Astronomy

Week 5: Stellar Evolution

Clip 13: Core Collapse



The Center Cannot Hold

- As gravitational crush increases, iron core collapses from size of Earth to a few km in 0.1s
- In core, $T_c \sim 3.5 \times 10^9 \, \mathrm{K}$ emits Υ rays leading to photodisintegration of heavy nuclei
- Outer layers fall inward at speeds up to 0.15c

- As core collapses electron degeneracy overcome
- Electrons forced into

$$p^+ + e^- \to n + \nu_e$$

- Left with a small, incredibly dense core that is mostly neutrons
- Does collapse stop?

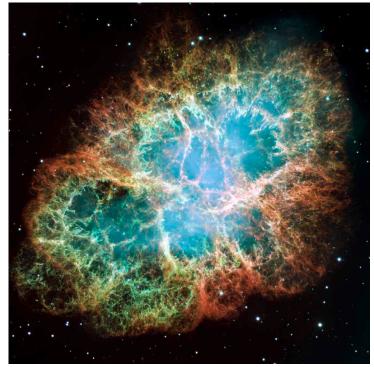


Boom!

- Within 0.25s core is neutrons with radius 20 km and super-nuclear density $\rho\sim 8\times 10^{17}{\rm kg/m^3}$
- Very little light can escape, energy carried off by neutrinos. Power emitted in these exceeds all known stars for 10 s
- At this density core collapse stops with bounce
- Colliding with infalling layers this triggers shock wave blowing outer star into space (96% of mass for $25M_{\odot}$ star)
- In compressed heated shock wave fusion to Fe and beyond via rprocess
- As ejecta thin light can escape. Luminosity reaches $3 imes 10^8 L_{\odot}$
- Energy released 10⁴⁷ J type-II supernova gravitational in origin

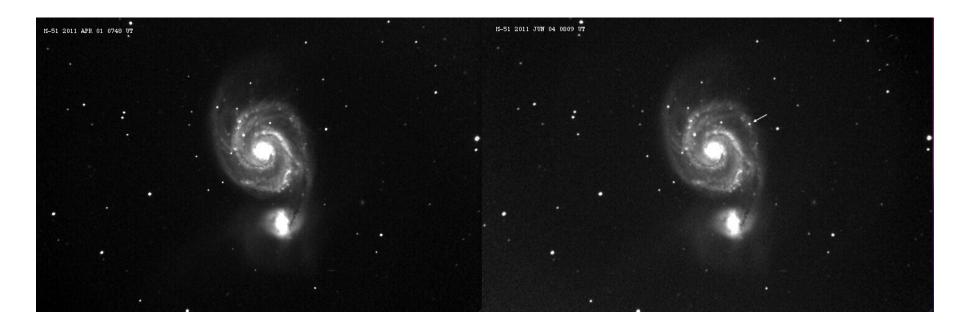
Seeing Them

- Sung dynasty history describes a supernova in 1054 whose remnant – Crab nebula in Taurus – is still visible (M1)
- Japanese, Arabic, Native American records concur
- Milky Way supernovae also in 1006, 1572, 1604. Estimated every 300 years but obscured by dust
- Many visible in other galaxies, currently some 20-30 bright ones





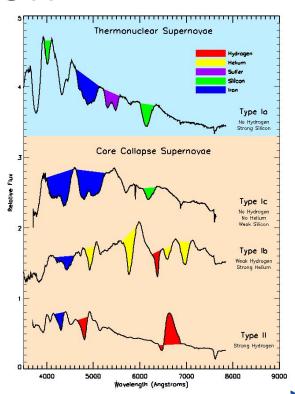
SN 2011dh





Classification

- SN classified by spectrum:
 - Ia: Strong Si no H He
 - Ib: Weak H Strong He
 - Ic: Weak Si no H He
 - II: Strong H
- Ia are nuclear explosion of WD
- II Ib Ic are gravitational core collapse with degrees of envelope loss



Credits

- M1: NASA, ESA, J. Hester and A. Loll (Arizona State University) http://hubblesite.org/newscenter/archive/releases/2005/37/image/a/
- SN 2011dh: P. Wiggins, Salt Lake Astronomical Society http://www.slas.us/gallery2/main.php?
 g2 itemId=4342&g2 imageViewsIndex=1
- SN Types: http://supernova.lbl.gov/~dnkasen/tutorial/graphics/sn_types.jpg

