Introductory Astronomy

Week 4: Stars

Clip 4: Neutrinos and a Surprise



How Do We Know?

- 1932) first
- Davis, Bahcall (1968): Detect the ν_e
- Pro: Penetrate Sun
- Con: Penetrate detector
- Flux at Earth: $10^{11} \frac{\nu_e}{\text{m}^2 \text{ s}}$

• Theory (Eddington, Bethe • Put a tank with $380 \,\mathrm{m}^3$ of **Chlorine** in Homestake Gold Mine

$$^{37}\text{Cl} + \nu_e \rightarrow ^{37}\text{Ar} + e$$

- Requires high-energy ν_e produced in other processes
- Expect one atom per six days



Where Are the Neutrinos?

- Flux Found is less than predictions
- Is Solar Model wrong?
- Is detector model wrong?
- Decided in 2001 by SNO: particle physics





More Particles, More Charges

Particle	Q	$N_{\rm e}$	N_{μ}	N_{τ}	Mass
p	1	0	0	0	935
n	0	0	0	0	938
e	-1	1	0	0	0.511
$ u_e$	0	1	0	0	?
μ	-1	0	1	0	106
$ u_{\mu}$	0	0	1	0	?
au	-1	0	0	1	1777
$ u_{\tau}$	0	0	0	1	,



So What?

- Neutrinos change spontaneously en route
- pp process produces ν_e
- When they arrive, 1/3 are ν_e
- This implies, in particular, that neutrinos are not massless although light.



Credits

 SNO: View of the SNO detector after installation of the bottom PMT panels, but before cabling. Photo courtesy of Ernest Orlando Lawrence Berkeley National Laboratory

http://www.sno.phy.queensu.ca/sno/images/
publicity_photos/index.html

