

Introductory Astronomy

Week 3: Solar System(s)

Clip 8: Terrestrial Planets

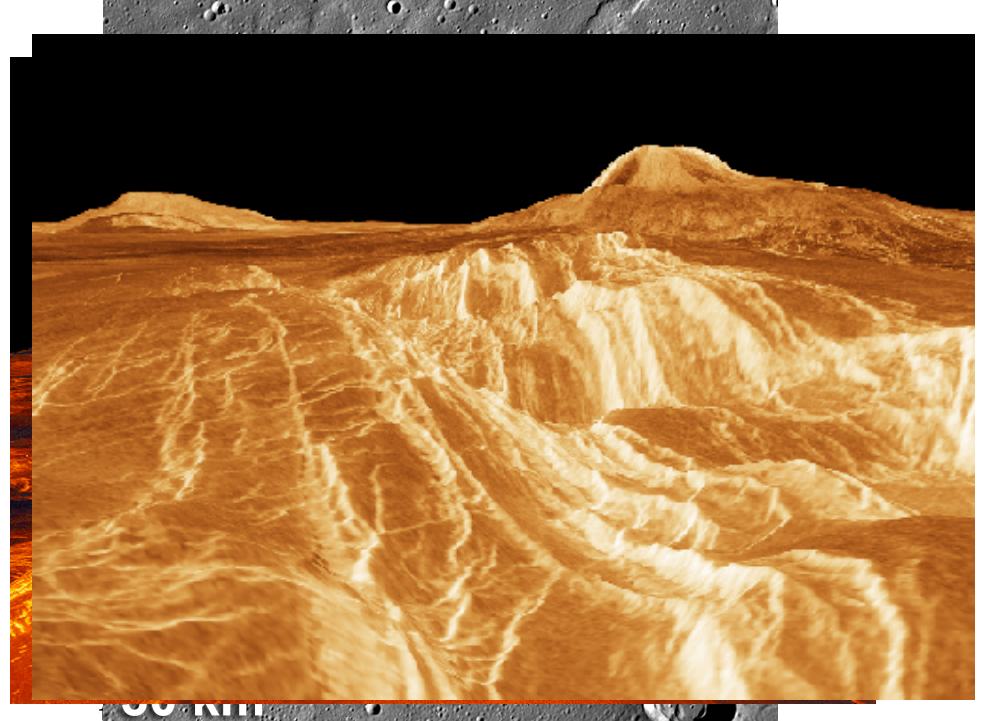
Spin and Orbit

- **Mercury**: close to Sun and eccentric: 2:3 tidal locking to 88d orbit tilted 7°
- **Venus**: 224d orbit tilted 3.4° ; 243d retrograde spin tilted 177.3°
- **Mars**: 687d orbit tilted 1.9° ; 24.6h spin tilted 25°

On the Surface

Tectonic Activity:

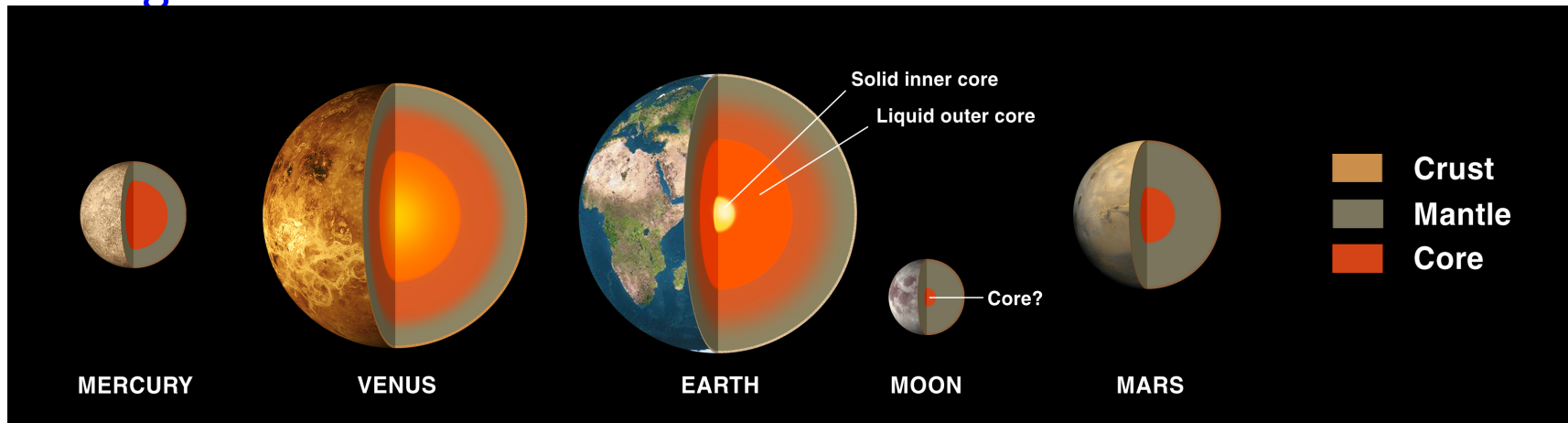
- **Mercury**: Features reflect cooling **compression** and **extension**
- **Venus**: Strong flaky crust and viscous mantle. **Tesserae**, **chasmate**. Global **Resurfacing 400Mya**
- **Mars**: **Two Plates?**
- Evidence of **volcanic** activity on **Mars** extinct





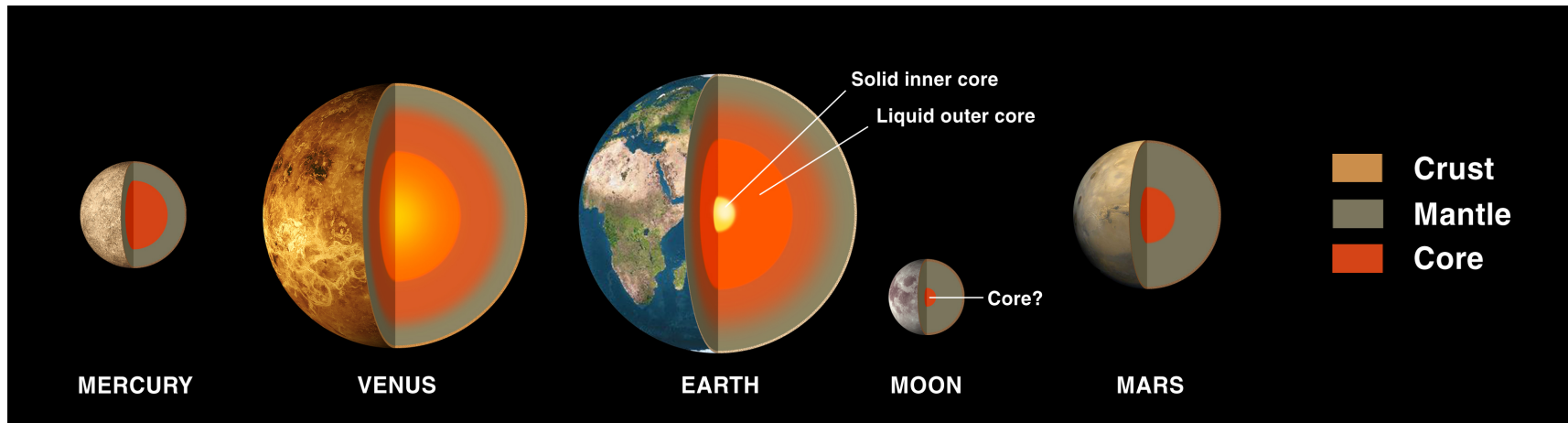
Inside

Mercury: Relatively huge core: uncompressed density 5300 kg/m^3 , at least partially liquid. Dipolar Magnetic Field



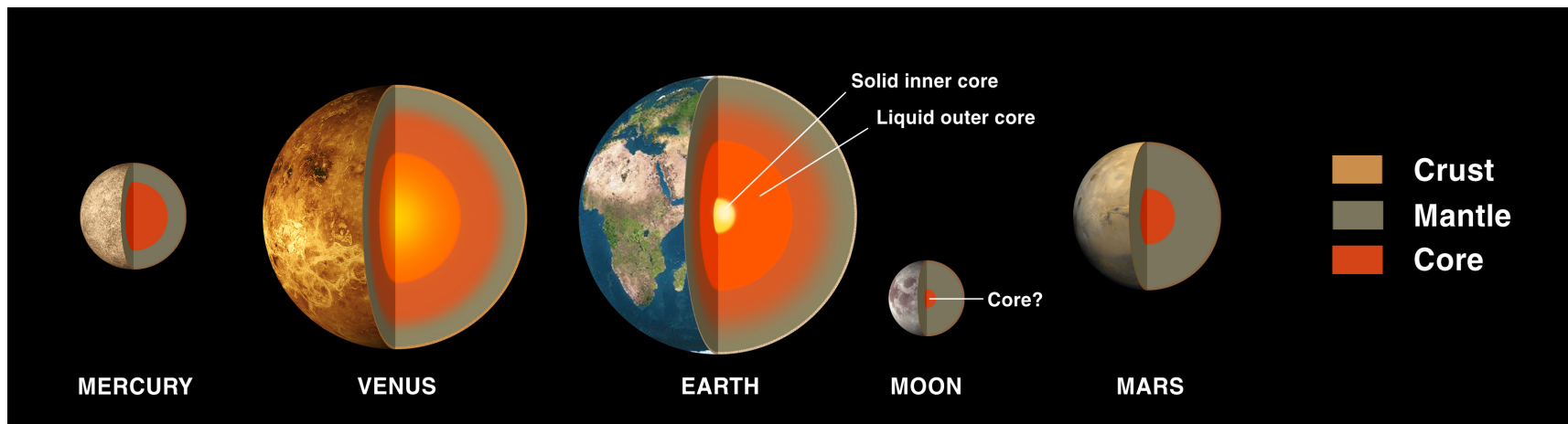
Inside

Venus: Presumed similar to Earth, density 5243 kg/m^3
No geodynamic **Magnetic Field**



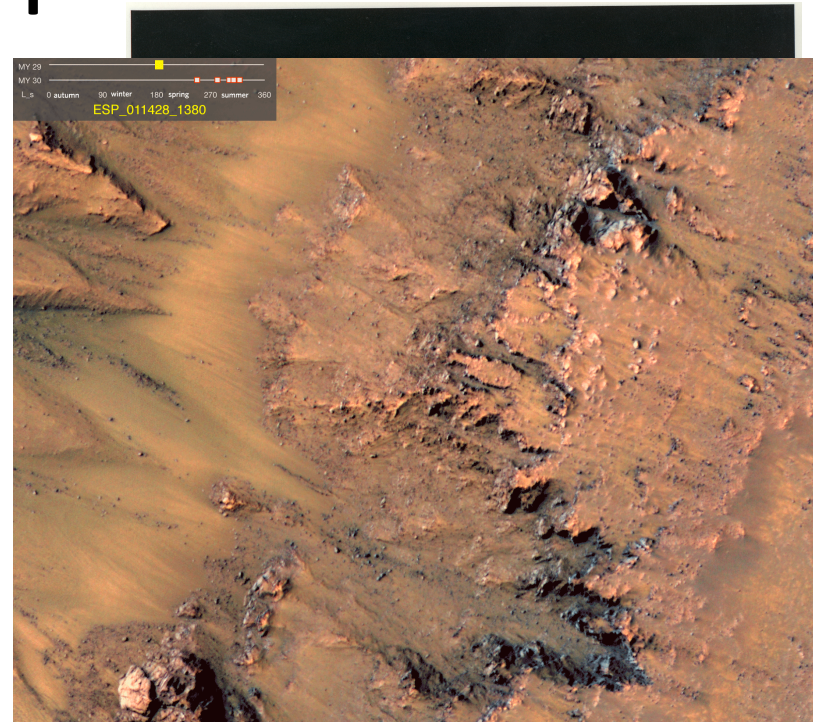
Inside

Mars: Small core, density 3934 kg/m^3 Mantle inactive. No **Magnetic Field** now but old **geodynamo**



Atmospheres

- **Mercury**: Not much. Albedo 0.142 $T=700\text{K}/100\text{K}$. Ice in shadows?
- **Venus**: Dense (93x Earth) CO_2 . SO_2 clouds Albedo 0.9 $T=730\text{K}$ No water.
- **Mars**: Tenuous (.005x Earth) CO_2 . Albedo 0.25 $T=308/130$ seasonal. Large polar ice caps. Water?



Why? A Comparative History

- Venus: Initially similar to Earth. CO_2 outgassed and water
- Evaporating water does not condense - dissociated and H_2 lost
- No oceans or tectonics CO_2 builds up
- Runaway greenhouse
- Mars: Initially similar to Earth. CO_2 outgassed and water
- Evaporating water condenses and rains, removing greenhouse gases
- No tectonics, gas trapped in rocks stays there
- Cooling traps water as permafrost if not lost to dissociation
- Weak magnetic field allows Solar Wind to strip remnant atmosphere
- Runaway Icehouse

Credits

- Mercury Orbit Animation: NASA JHU/APL
<http://sciencenetlinks.com/interactives/messenger/or/OrbitRotation.html>
- Mercury Surface: NASA/Johns Hopkins University Applied Physics Laboratory/Carnegie Institution of Washington
http://messenger.jhuapl.edu/gallery/sciencePhotos/pics/Watters_Press_photo_1.jpeg
- Venus Surface Image: NASA/JPL <http://www2.jpl.nasa.gov/magellan/image13.html>
- Mars: NASA http://solarsystem.nasa.gov/multimedia/display.cfm?Category=Planets&IM_ID=2050
- Mars Topography: NASA <http://photojournal.jpl.nasa.gov/jpeg/PIA02031.jpg>
- Internal Structure: NASA http://solarsystem.nasa.gov/multimedia/display.cfm?IM_ID=168
- HD 17255 Animation: NASA/JPL-Caltech
http://www.nasa.gov/mission_pages/spitzer/multimedia/spitzer-20090810.html
- Venus Clouds: NASA http://nssdc.gsfc.nasa.gov/image/planetary/venus/pvo_uv_790205.jpg
- Mars Icecap: Philip James (University of Toledo), Steven Lee (University of Colorado),
NASA http://nssdc.gsfc.nasa.gov/image/planetary/mars/hst_mars060.jpg
- Water on Mars: NASA/JPL-Caltech/Univ. of Arizona
http://www.nasa.gov/mission_pages/MRO/news/mro20110804.html