

Galaxy Scaling Relations Part I

Galaxy Scaling Laws

- When correlated, global properties of galaxies tend to do so as power-laws; thus "scaling laws"
- They provide a quantitative means of examining physical properties of galaxies and their systematics
- They reflect the internal physics of galaxies, and are a product of the formative and evolutionary histories
 - Thus, they could be (and are) different for different galaxy families
 - We can use them as a fossil evidence of galaxy formation
- When expressed as correlations between distance-dependent and distance-independent quantities, they can be used to measure relative distances of galaxies and peculiar velocities: thus, it is really important to understand their intrinsic limitations of accuracy, e.g., environmental dependences

The Tully-Fisher Relation

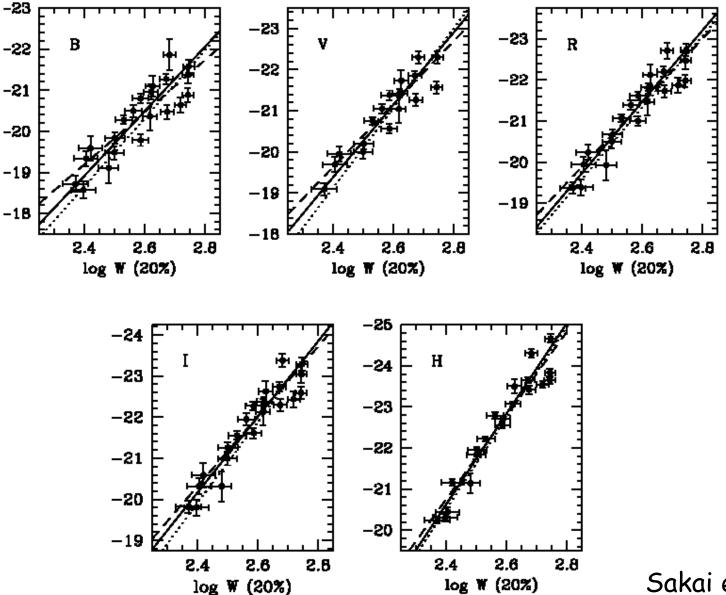
• A well-defined luminosity vs. rotational speed (often measured as a H I 21 cm line width) relation for spirals:

 $L \sim V_{rot}^{\gamma}$, $\gamma \approx 4$, varies with wavelength

Or: $M = b \log (W) + c$, where:

- *M* is the absolute magnitude
- W is the Doppler broadened line width, typically measured using the HI 21cm line, corrected for inclination $W_{true} = W_{obs}/sin(i)$
- Both the slope b and the zero-point c can be measured from a set of nearby spiral galaxies with well-known distances
- The slope b can be also measured from any set of galaxies with roughly the same distance - e.g., galaxies in a cluster even if that distance is not known
- Scatter is $\sim 10-20\%$ at best, better in the redder bands

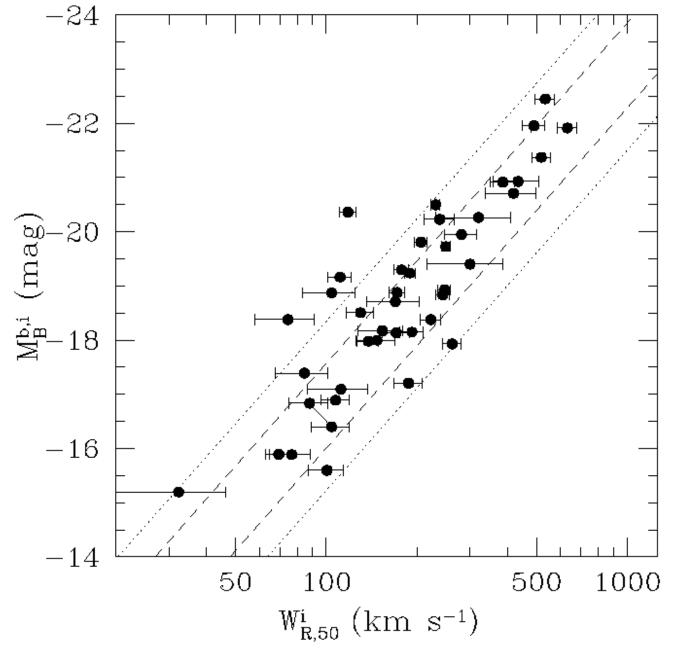
Tully-Fisher Relation in Diferent Bands



Sakai et al 1999

Why is the TFR So Remarkable?

- Because it connects a property of the dark halo the maximum circular speed with the product of the net integrated star formation history, i.e., the luminosity of the disk
- Halo-regulated galaxy formation/evolution?
- The scatter is remarkably low even though the conditions for this to happen are known not to be satisfied
- There is some important feedback mechanism involved, which we do not understand yet
- Thus, the TFR offers some important insights into the physics of disk galaxy formation

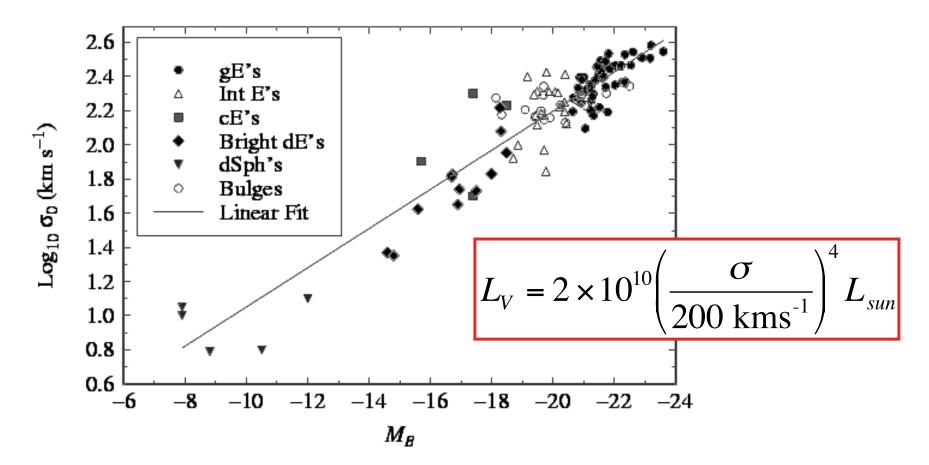


Low surface brightness galaxies follow the <u>same</u> TF law as the regular spirals: so it is really relating the baryonic mass to the dark halo

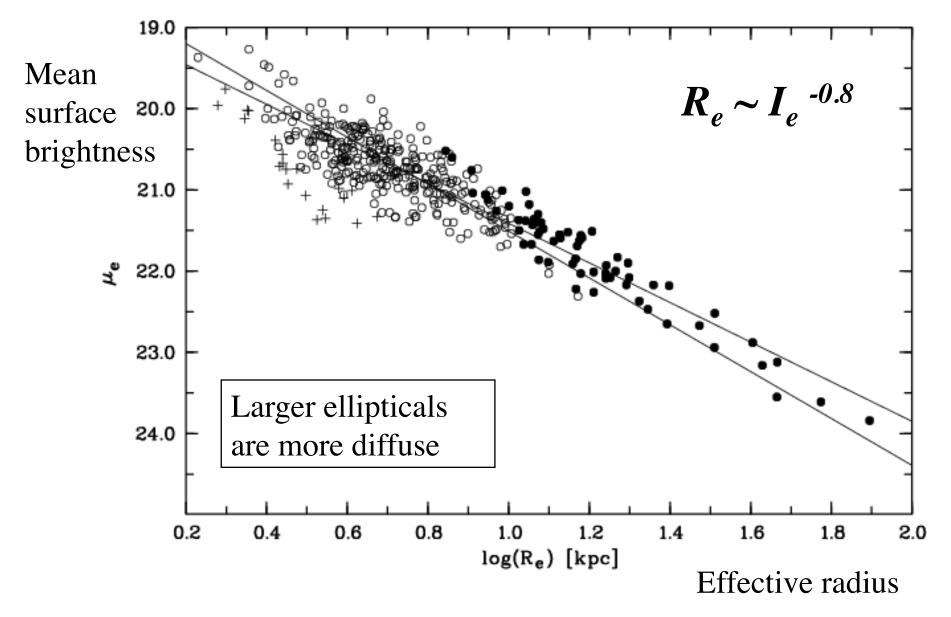
Zwaan et al. 1995

The Faber-Jackson Relation for Ellipticals

Analog of the Tully-Fisher relation for spirals, but instead of the peak rotation speed V_{max} , measure the velocity dispersion. This is correlated with the total luminosity:



The Kormendy Relation for Ellipticals



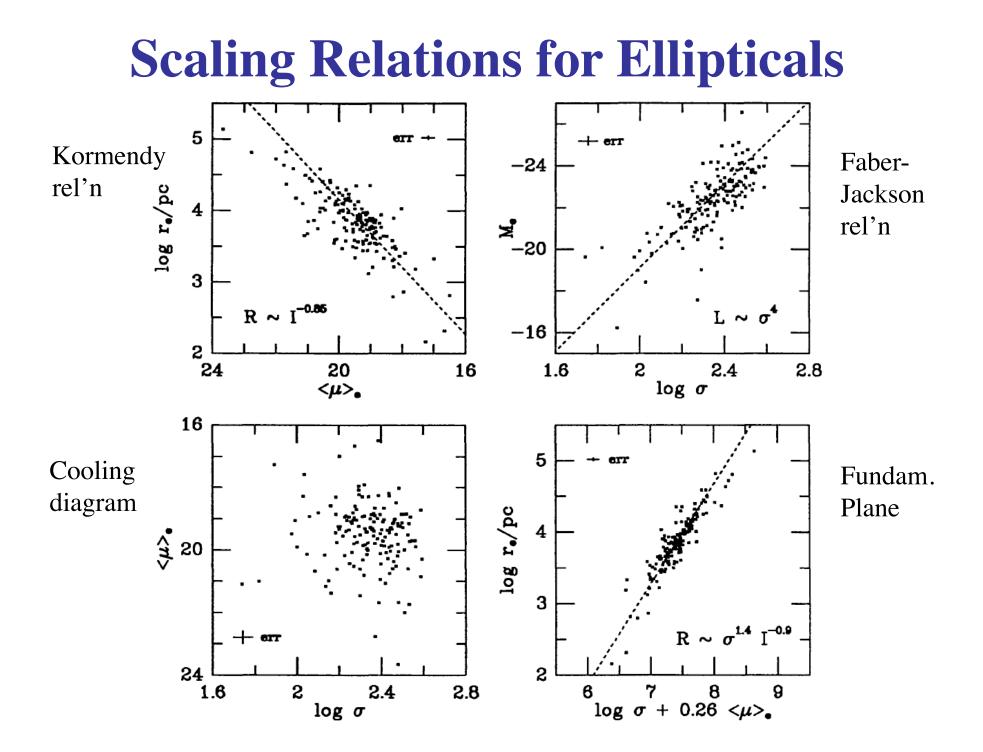
Can We Learn Something About the Formation of Ellipticals From the Kormendy Relation?

From the Virial Theorem, $m\sigma^2 \sim GmM/R$ Thus, the dynamical mass scales as $M \sim R\sigma^2$ Luminosity $L \sim I R^2$, where I is the mean surface brightness Assuming $(M/L) = const., M \sim I R^2 \sim R\sigma^2$ and $I R \sim \sigma^2$

Now, if ellipticals form via dissipationless merging, the kinetic energy per unit mass ~ σ^2 ~ *const.*, and thus we would predict the scaling to be $R \sim I^{-1}$

If, on the other hand, ellipticals form via dissipative collapse, then M = const., surface brightness $I \sim M R^{-2}$, and thus we would predict the scaling to be $R \sim I^{-0.5}$

The observed scaling is $R \sim I^{-0.8}$. Thus, *both* dissipative collapse and dissipationless merging probably play a role



Deriving the Scaling Relations

Start with the Virial Theorem:

$$\frac{GM}{\langle R \rangle} = k_E \frac{\langle V^2 \rangle}{2}$$

Now relate the observable values of R, V (or σ), L, etc., to their "true" mean 3-dim. values by simple scalings:

$$R = k_R \langle R \rangle \qquad V^2 = k_V \langle V^2 \rangle \qquad L = k_L I R^2$$

One can then derive the "virial" versions of the FP and the TFR:

Where the "structure" coefficients are:

$$R = K_{SR} V^2 I^{-1} (M/L)^{-1}$$
$$L = K_{SL} V^4 I^{-1} (M/L)^{-2}$$

$$K_{SR} = \frac{k_E}{2Gk_Rk_Lk_V}$$
$$K_{SL} = \frac{k_E^2}{4G^2k_R^2k_Lk_V^2}$$

Deviations of the observed relations from these scalings must indicate that either some k's and/or the (M/L) are changing

Galaxy Scaling Relations Will Continue in Part II

