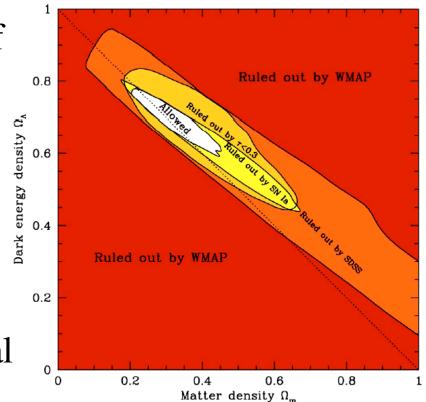
The Dark Energy

The Dark Energy

- The **dominant component** of the observed matter/energy density: $\Omega_{0,DE} \approx 0.7$
- Causes the accelerated expansion of the universe
- May affect the growth of density perturbations
- Effective only at cosmological distances



- Its physical nature is as yet *unknown*; this may be the biggest outstanding problem in physics today
- *Cosmological constant* is just one special case; a more general possibility is called *quintessence*

Cosmological Constant

It classically appears as an integration constant in Einstein Eqs:

$$G_{\mu\nu} = 8\pi G T_{\mu\nu} - \Lambda g_{\mu\nu}$$

where $\Lambda = 1/L^2, \ \rho = \Lambda/4\pi G$

Note that it does not have a preferred value or explanation in GR (and thus cannot be declared a priori to be zero); it becomes a new constant of nature, and joins G in defining the gravity

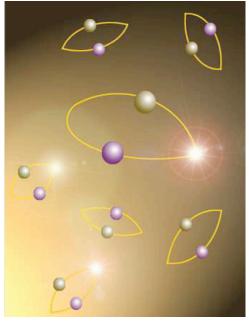
In the Friedmann Eq., it acts as an elastic force, proportional to the distance; so you can think of it as the *elasticity of the physical vacuum*.

By the standard convention, $\Lambda > 0$ (positive energy density) corresponds to a repelling force, and vice versa.

Cosmological Constant as a Quantum Field Phenomenon

- Proposed by Yakov Zeldovich (1967)
- A modern view of the physical vacuum is that it is not really empty - it is filled with virtual particle-antiparticle pairs, which annihilate within $\Delta t < \hbar/mc^2$, and their fluctuations give rise to a net energy density - a ground(?) state of the physical vacuum
- This is essentially the same mechanism
- Brandsred any the origin of the inflations vacuum energy density, we need a quantum theory of gravity, which we don't have yet
- Nevertheless, eager minds do try ...





The Worst Scientific Prediction Ever

• A "natural" Planck system of units expresses everything as combination of fundamental physical constants; the Planck density is:

$$\rho_{Planck} = c^{5} / (\hbar G^{2}) = 5.15 \times 10^{+93} \text{ g cm}^{-3}$$

• The observed value is:

 $\rho_{vac} = \Omega_{vac} \ \rho_{crit} \approx 6.5 \times 10^{-30} \text{ g cm}^{-3}$ Ooops! Off by 123 orders of magnitude ...

- This is modestly called "the fine-tuning problem" (because it requires a cancellation to 1 part in 10¹²³)
- The other "natural" value is zero
- So, lacking a proper theory, physicists just declared the cosmological constant to be zero, and went on...

Cosmological Constant or Quintessence?

- **Cosmological constant:** energy density constant in time and spatially uniform
 - Corresponds to the energy density of the physical vacuum
 - A coincidence problem: why is $\Omega_{\Lambda} \sim \Omega_{m}$ just now?
- **Quintessence:** time dependent and possibly spatially inhomogeneous; e.g. scalar field rolling down a potential
- Both can be described in the equation of state formalism:

$$P = w \rho$$

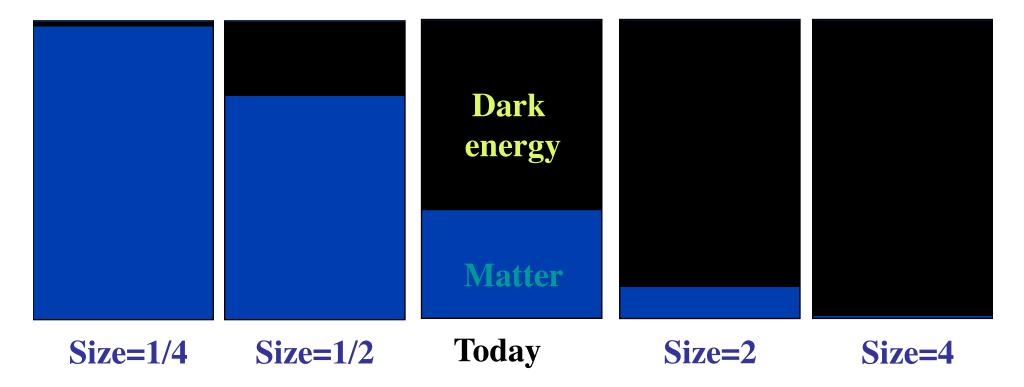
 $\rho \sim R^{-3(w+1)}$

Cosmological constant: $w = const. = -1, \rho = const.$

Quintessence: *w* can have other values and change in time

The Cosmic Coincidence Problem

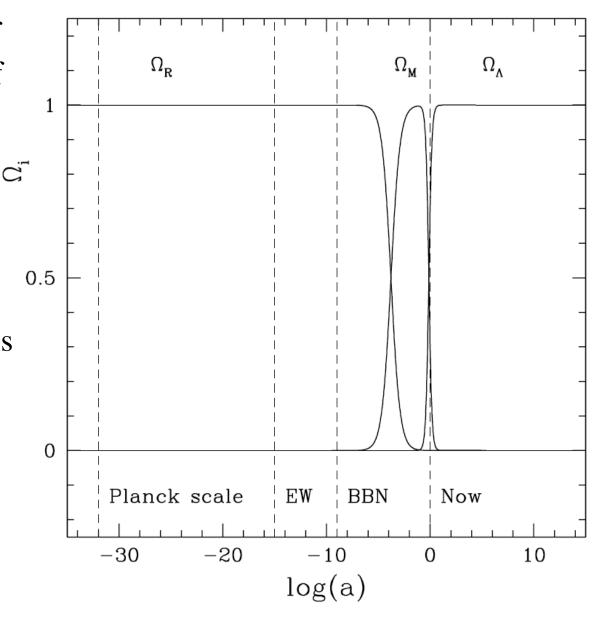
If the dark energy is really due to a cosmological constant, its density does not change in time, whereas the matter density does - and they just happen to be comparable today! Seems un-natural ...



The Cosmological Coincidence Problem

The time dependence of the density parameter of various mass/energy density components:

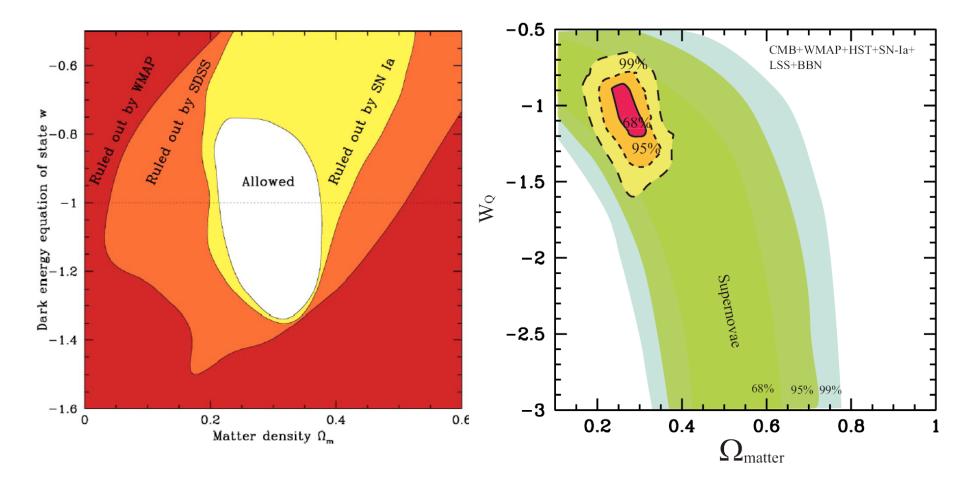
We seem to live in a special era, when the vacuum energy density is just starting to dominate the dynamics of the universe ...



Physical Origins of the Dark Energy

- ... are completely unknown at this time, and not for the lack of trying: there are literally thousands of papers about it, and more being published every day
- Many of the proposed models are based on one of the following:
 - Decay of some scalar field, similar to the inflation mechanism
 - Modified theories of gravity
 - Holographic models, connecting the vacuum energy density to the area of the event horizon and thermodynamics
 - Landscape or multiverse models that postulate the existence of $\sim 10^{500}$ separate universes, with different (random) values of the physical constants, Λ included
 - Models connecting DM and DE ... *etc., etc.*
- One measurement that might help eliminate some possibilities is a possible deviation (evolution) of the EOS parameter *w*

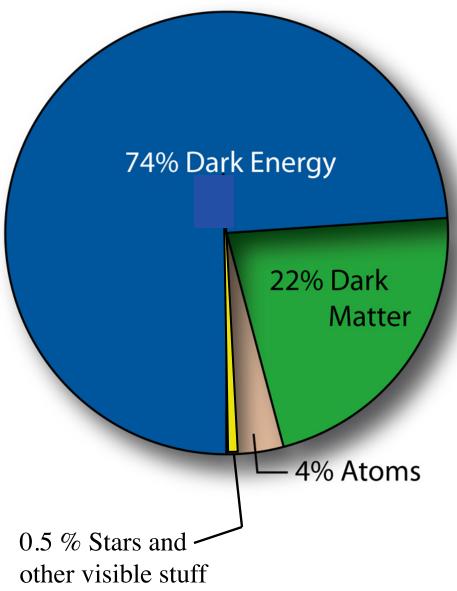
Observational Constraints on W Strongly favor values of $w \sim -1$, i.e., *cosmological constant* like dark energy. Some models can be excluded, but there is still some room for $\rho_{vac} \neq const$. models



Observational Constraints on *W* gentle acceleration -0.6 BAO -0.7 all SN 1a -0.8 dark energy "stiffness" -0.9 Cosmological -1.0 constant WMAP -1.1 **Big Rip** -1.2 -1.3 clusters +WMAP -1.4 -1.5 0.85 0.60 0.65 0.70 0.75 0.80

dark energy density

Contents of the Universe: Summary



- $\Omega_0 = 1.00 \pm 0.02$
- $\Omega_m \approx 0.27 \pm 20\%$
 - $-\Omega_b \approx 0.045 \pm 10\%$ \circ Includes $\Omega_{visible} \approx 0.005$

$$-\Omega_{non-b}\approx 0.22$$

 \circ Includes $\Omega_{v} < 0.005$

$$-\Omega_{CMBR} \approx 0.0001$$

•
$$\Omega_{de} \approx 0.73 \pm 10\%$$

• The physical nature of the DE is currently completely unknown



Structure Formation: