Tests for the Expansion of the Universe







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- Tolman surface brightness (SB) test
 - In a stationary, Euclidean universe SB = const.
 - In an expanding universe, $SB \sim (1+z)^{-4}$
 - In a "tired light" model, $SB \sim (1+z)^{-1}$
- Time dilation of Supernova light curves
 - Time stretches by a factor of (1+v/c) = (1+z)
- The match between the energy density and *T*⁴ for the blackbody and the CMBR
 - For a blackbody, energy density $u \sim T^4$
 - In an expanding universe, for photons, energy density is $u \sim (1+z)^4$, and since $T \sim 1/\lambda \sim (1+z)$, $u \sim T^4$

The Tolman Test

Surface brightness is flux per unit solid angle: $B = \frac{f}{d\omega}$

This is the same as the luminosity per unit area, at some distance D. In cosmology, $B = \frac{L}{D_{\perp}^2} \frac{D_A^2}{dl^2}$

In a stationary, Euclidean case, $D = D_L = D_A$, so the distances cancel, and SB = const. But in an expanding universe, $D_L = D (1+z)$, and $D_A = D / (1+z)$, so:

$$B = \frac{L}{dl^2} \frac{D_A^2}{D_L^2} = \frac{L}{dl^2} (1+z)^{-4}$$

Note that this is independent of cosmology!

Performing the The Tolman Test

We need a standard (constant) unit of surface brightness = luminosity/area, to observe at a range of redshifts (a "standard fuzz"?)

A good choice is the intercept of surface brightness scaling relations for elliptical galaxies in clusters



The Tolman Test Results



Surface brightness intercept of the Fundamental Plane correlation, for elliptical galaxies in clusters out to z ~ 0.6. It assumes a reasonable galaxy evolution model correction.

⁽from Pahre et al.)

Time Dilation of Supernova Lightcurves



Time Dilation of Supernova Lightcurves



Normalized lightcurve width increases linearly with (1+z), as expected in an expanding universe

After correcting for the expansion, the stretch is constant

(Goldhaber et al.)

