# Ay 21 - Galaxies and Cosmology Prof. S. G. Djorgovski



# **Cosmology\* as a Science**

- A study of the universe as a whole, its global geometry, dynamics, history, fate, and its major constituents galaxies and large-scale structures, their formation and evolution
- A basic assumption: the physical laws are the same at all times and everywhere
  - Some aspects of this are testable
  - But a new and unexpected physics can show up, e.g., dark matter, dark energy
- Only one object of study, and all we can do is look at the surface of the past light cone
- Observations tend to be difficult, and subject to biases and selection effects

<sup>\*</sup> From Greek *kosmos* = order; see also *cosmetology* ...

#### The Evolution of the Cosmological Thought

... From magical and arbitrary to rational and scientific Folklore to theology to philosophy to physics

... Away from anthropocentric/anthropomorphic

The Copernican revolution

- ... From final and static to evolving and open-ended The Darwinian revolution
- ... From absolute certainty to an ever expanding sphere of knowledge and a boundary of unknown

Cosmology today is a branch of physics



# **Dust Off Your Astronomical Units!**

- Distance:
  - Astronomical unit: the distance from the Earth to the Sun, 1 au =  $1.496 \times 10^{13}$  cm
  - Light year: c  $\times 1$  yr, 1 ly = 9.463  $\times 10^{17}$  cm
  - Parsec: the distance from which 1 au subtends an angle of 1 arcsec, 1 pc =  $3.086 \times 10^{18}$  cm = 3.26 ly = 206,264.8 au
- Mass and Luminosity:
  - Solar mass:  $1 M_{\odot} = 1.989 \times 10^{33} g$
  - Solar luminosity:  $1 L_{\odot} = 3.826 \times 10^{33} \text{ erg/s}$

### Fluxes

Real detectors are sensitive over a finite range of  $\lambda$  (or  $\nu$ ). Fluxes are always measured over some finite bandpass.

Total energy flux:  

$$F = \int F_v(v) dv$$
 Integral of  $f_v$  over  
all frequencies  
Units: erg s<sup>-1</sup> cm<sup>-2</sup> Hz<sup>-1</sup>

A standard unit for specific flux (initially in radio, but now more common): 1 Jansky  $(Jy) = 10^{-23} \text{ erg s}^{-1} \text{ cm}^{-2} \text{ Hz}^{-1}$ 

I Jansky 
$$(Jy) = 10^{25} \text{ erg s}^{-1} \text{ cm}^{-2} \text{ Hz}^{-1}$$

 $f_v$  is often called the *flux density* - to get the *power*, one integrates it over the bandwith, and multiplies by the area

### **Fluxes and Magnitudes**

For historical reasons, fluxes in the optical and IR are measured in magnitudes:  $m = -2.5 \log_{10} F + \text{constant}$ 



If the flux is integrated over the entire spectrum, then *m* is the *bolometric magnitude*.

# **Magnitude Zero Points**



 $AB_{\rm v} = -2.5 \log f_{\rm v} \,[{\rm cgs}] - 48.60$ 

# **Using Magnitudes**

Consider two stars, one of which is a hundred times fainter than the other in some waveband (say V).

$$m_{1} = -2.5 \log F_{1} + \text{constant}$$
  

$$m_{2} = -2.5 \log(0.01F_{1}) + \text{constant}$$
  

$$= -2.5 \log(0.01) - 2.5 \log F_{1} + \text{constant}$$
  

$$= 5 - 2.5 \log F_{1} + \text{constant}$$
  

$$= 5 + m_{1}$$

Source that is 100 times **fainter** in flux is five magnitudes fainter (**larger** number).

Faintest objects detectable with *HST* have magnitudes of ~ 29 in R/I bands. The sun has  $m_V = -26.75$  mag

# **Apparent vs. Absolute Magnitudes**

The absolute magnitude is defined as the apparent mag. a source would have if it were at a distance of 10 pc:

$$M = m + 5 - 5 \log d/\mathrm{pc}$$

It is a measure of the **luminosity** in some waveband.

For Sun: 
$$M_{\odot B} = 5.47, M_{\odot V} = 4.82, M_{\odot bol} = 4.74$$

Difference between the apparent magnitude *m* and the absolute magnitude *M* (any band) is a *measure of the distance* to the source

$$m - M = 5\log_{10}\left(\frac{d}{10 \text{ pc}}\right)$$

**Distance modulus** 

The Olbers Paradox (1826): Why is the sky dark at night?

Assume:

- The universe is infinite
- The universe is filled uniformly with stars
- The universe is eternal and unchanging

#### Deriving the Olbers Paradox:

Total flux we receive from a star  $\propto 1/d^2$ 

Angular area of a star  $\propto 1/d^2$ 

Surface brightness (light per unit area on sky)

= total flux / angular area = constant

If every line of sight ends on a star, the sky should be as bright as the sun!



The equivalent paradox applies to gravity - infinite tidal forces?!

# **Solving the Olbers Paradox**

- Opaque universe? *No, the energy is still there*
- Redshift? Not enough
- Hierarchical universe? *Possible, but not true*

#### The actual explanation:

- The universe has a finite extent, age, and energy content: there isn't enough radiative energy in it to make the sky as bright as the surface of a star
  - In fact, the sky is ablaze but the temperature of the radiation is only 2.7° K (CMBR)
  - All starlight ever emitted amounts only to a few percent of the CMBR energy density

#### Next: Early History of Cosmology: From ~ 18<sup>th</sup> Century Until ~ 1930's

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